

FALL3D-6.2

User Guide

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1 Introduction

FALL3D-6.2 is a 3-D time-dependent Eulerian model for the transport and deposition of volcanic ash. The model solves the advection-diffusion-sedimentation (ADS) equation on a structured terrain-following grid using a second-order Finite Differences (FD) explicit scheme. Different parameterizations for the eddy diffusivity tensor and for the particle terminal settling velocity can be chosen. The code, written in FORTRAN-90, is available for Unix/Linux/Mac X Operating Systems (OS). A set of pre- and post-process utility programs and OS-dependent scripts to launch them are also included in the FALL3D-6.2 distribution package. Although the model has been designed to forecast volcanic ash concentration in the atmosphere and ash loading at the ground, it can also be used to model the transport of other kinds of airborne solid particles. The model inputs are meteorological data, topography, grain-size distribution, shape and density of particles, and mass rate of particle injected into the atmosphere. The FALL3D-6.2 model can be used as a tool for short-term ash deposition forecasting and for volcanic fallout hazard assessment. FALL3D-6.2 is available in two versions called PUB (Public) version and PROF (Professional) version.

More information on <http://www.bsc.es/projects/earthscience/fall3d/> and <http://datasim.ov.ingv.it/Fall3d.html>

2 Ash transport model

In this section we briefly describe the governing equations and the main assumptions used in FALL3D-6.2. For further details see Costa et al. (2006); Folch et al. (2009).

2.1 Governing equations

The main factors controlling atmospheric transport of ash are wind advection, turbulent diffusion, and gravitational settling of particles. Neglecting particle-particle interaction effects (collisions, aggregation, *etc.*), the Eulerian form of the continuity equation written in a generalized coordinate system (X, Y, Z) is (Byun and Schere, 2006; Costa et al., 2006):

$$\begin{aligned} \frac{\partial C}{\partial t} + V_X \frac{\partial C}{\partial X} + V_Y \frac{\partial C}{\partial Y} + (V_Z - V_{sj}) \frac{\partial C}{\partial Z} = -C \nabla \cdot \mathbf{V} + C \frac{\partial V_{sj}}{\partial Z} \\ + \frac{\partial}{\partial X} \left(\rho_* K_X \frac{\partial C / \rho_*}{\partial X} \right) + \frac{\partial}{\partial Y} \left(\rho_* K_Y \frac{\partial C / \rho_*}{\partial Y} \right) + \frac{\partial}{\partial Z} \left(\rho_* K_Z \frac{\partial C / \rho_*}{\partial Z} \right) + S_* \end{aligned} \quad (1)$$

where C is the transformed concentration, $\mathbf{V} = (V_X, V_Y, V_Z)$ is the transformed wind speed, K_X, K_Y and K_Z are the diagonal terms of the transformed eddy diffusivity tensor, ρ_* is the transformed atmospheric density, and S_* is the transformed source term. FALL3D-6.2 solves Eq. (1) for each particle class j using a curvilinear terrain-following coordinate system ($X = mx, Y = my, z \rightarrow Z$), where m is the map scale factor and $Z = z - h(x, y)$, with $h(x, y)$ denoting the topographic elevation, and (x, y, z) are the Cartesian coordinates. The scaling factors for this particular transformation are given in Table 1 (Byun and Schere, 2006). The generic particle class j is defined by a triplet of values characterizing each particle (d_p, ρ_p, F_p) , that are, respectively, diameter, density, and a shape factor. For d_p we use the equivalent diameter d , which is the diameter of a sphere of equivalent volume. For the shape factor F_p we choose the sphericity ψ , which is the ratio of the surface area of a sphere with diameter d to the surface area of the particle. In our approximation, each triplet (d, ρ_p, ψ) is sufficient to define the settling velocity. Effect of Earth's curvature are considered when the lat-lon coordinate system is used through the Jacobian of the transformation.

2.2 Eddy Diffusivity Tensor

In FALL3D-6.2 only the diagonal components of the Eddy Diffusivity Tensor, *i.e.* the vertical K_z and the horizontal $K_h = K_x = K_y$ components, are considered.

The available choices for describing the vertical component K_z are:

1. Option CONSTANT, *i.e.* $K_z = constant$, where the constant value is assigned by the user;

Parameter	Scaling
Coordinates	$X = mx; \quad Y = my; \quad Z = z - h(x, y)$
Horizontal Velocities	$V_X = mv_x; \quad V_Y = mv_y$
Vertical velocity	$(V_Z - V_{Sj}) = J^{-1} \left[(v_z - v_{sj}) - m \left(v_x \frac{\partial h}{\partial x} + v_y \frac{\partial h}{\partial y} \right) \right]$
Diffusion Coefficients	$K_X = K_x; \quad K_Y = K_y; \quad K_Z \simeq K_z J^{-2}$
Concentration	$C = cJ/m^2$
Density	$\rho_* = \rho J/m^2$
Source Term	$S_* = SJ/m^2$

Table 1: Scaling factors for a terrain-following coordinate system ($x = mX$, $y = mY$, $z \rightarrow Z$). (x, y, z) are the Cartesian coordinates, m the map scale factor (for the UTM coordinate system $m = 1$) and J is the determinant of the Jacobian of the coordinate system transformation.

2. Option **SIMILARITY**. In this case, inside the Atmospheric Boundary Layer (ABL), FALL3D-6.2 evaluates K_z as:

$$K_z = \begin{cases} \kappa u_* z \left(1 - \frac{z}{h}\right) \left(1 + 9.2 \frac{h}{L} \frac{z}{h}\right)^{-1} & h/L \geq 0 \quad \text{stable} \\ \kappa u_* z \left(1 - \frac{z}{h}\right) \left(1 - 13 \frac{h}{L} \frac{z}{h}\right)^{1/2} & h/L \leq 0 \quad \text{unstable} \end{cases} \quad (2)$$

where κ is the von Karman constant ($\kappa = 0.4$), u_* is the friction velocity, h is the ABL height, and L is the Monin-Obukhov length (see Costa et al., 2006). The expression above comes from an extension of the Monin-Obukhov similarity theory to the entire ABL (Ulke, 2000). On the other hand, above the ABL ($z/h > 1$), K_z is considered a function of the local vertical wind gradient, a characteristic length scale l_c , and a stability function F_c which depends on the Richardson number Ri :

$$K_z = l_c^2 \left| \frac{\partial U}{\partial z} \right| F_c(Ri) \quad (3)$$

where $U = \sqrt{y_x^2 + u_y^2}$. For l_c and F_c , FALL3D-6.2 adopts the relationship used by the CAM-3.0 model (Collins et al., 2004):

$$l_c = \left(\frac{1}{\kappa z} + \frac{1}{\lambda_c} \right)^{-1} \quad (4)$$

$$F_c(Ri) = \begin{cases} \frac{1}{1 + 10Ri(1 + 8Ri)} & \text{stable} \quad (Ri > 0) \\ \sqrt{1 - 18Ri} & \text{unstable} \quad (Ri < 0) \end{cases} \quad (5)$$

where λ_c is the so-called asymptotic length scale ($\lambda_c \approx 30\text{m}$).

The available choices for describing the horizontal component $K_h = K_x = K_y$ are:

1. Option **CONSTANT**, *i.e.* $K_h = \text{constant}$, where the constant value is assigned by the user;
2. Option **RAMS**. In this case, a large eddy parameterization as that used by the RAMS model (Pielke et al., 1992) can be used for evaluating K_h :

$$K_h = Pr_t \max \left(k_m; (C_S \Delta)^2 \sqrt{\left(\frac{\partial v_x}{\partial y} + \frac{\partial v_y}{\partial x} \right)^2 + 2 \left[\left(\frac{\partial v_x}{\partial x} \right)^2 + \left(\frac{\partial v_y}{\partial y} \right)^2 \right]} \right) \quad (6)$$

where Pr_t is the turbulent Prandtl number (typically $Pr_t \approx 1$), $k_m = 0.075\Delta^{4/3}$, $\Delta = \sqrt{\Delta x \Delta y}$, Δx and Δy are the horizontal grid spacings, and C_S is a constant ranging from 0.135 to 0.32.

3. Option **CMAQ**. In this case, the horizontal diffusion is evaluated as in the CMAQ model (Byun and Schere, 2006):

$$\frac{1}{K_h} = \frac{1}{K_{ht}} + \frac{1}{K_{hn}} \quad (7)$$

where:

$$K_{ht} = \alpha^2 \Delta x \Delta y \sqrt{\left(\frac{\partial v_x}{\partial x} - \frac{\partial v_y}{\partial y}\right)^2 + \left(\frac{\partial v_y}{\partial x} + \frac{\partial v_x}{\partial y}\right)^2} \quad (8)$$

$$K_{hn} = K_{hf} \left(\frac{\Delta x_f \Delta y_f}{\Delta x \Delta y}\right) \quad (9)$$

where $\alpha = 0.28$ and the values of K_{hf} and $\Delta x_f = \Delta y_f$ depend on the algorithm. Using this parameterization, for a large grid size the effect of the transportive dispersion is minimized, whereas for a small grid size the numerical diffusion term is reduced (Byun and Schere, 2006). Thanks to the heuristic relationship (7), the smaller between K_{ht} and K_{hn} dominates. In our case we set $K_{hf} = 8000 \text{ m}^2\text{s}^{-1}$ for $\Delta x_f = \Delta y_f = 4 \text{ km}$ and also a minimum value for K_h equal to $k_m = 0.075 \Delta^{4/3}$ was imposed.

2.3 Settling velocity models

There are several semi-empirical parameterizations for the particle settling velocity v_s if one assumes that particles settle down at their terminal velocity:

$$v_s = \sqrt{\frac{4g(\rho_p - \rho_a)d}{3C_d\rho_a}} \quad (10)$$

where ρ_a and ρ_p denote air and particle density, respectively, d is the particle equivalent diameter, and C_d is the drag coefficient. C_d depends on the Reynolds number, $Re = dv_s/\nu_a$ ($\nu_a = \mu_a/\rho_a$ is the kinematic viscosity of air, μ_a the dynamic viscosity). In FALL3D-6.2 different options are possible for estimating settling velocity, such as:

1. **ARASTOPOUR** model (Arastoopour et al., 1982):

$$C_d = \begin{cases} \frac{24}{Re}(1 + 0.15Re^{0.687}) & Re \leq 988.947 \\ 0.44 & Re > 988.947 \end{cases} \quad (11)$$

valid for spherical particles only.

2. **GANSER** model (Ganser, 1993):

$$C_d = \frac{24}{ReK_1} \left\{ 1 + 0.1118 (Re K_1 K_2)^{0.6567} \right\} + \frac{0.4305K_2}{1 + \frac{3305}{Re K_1 K_2}} \quad (12)$$

where $K_1 = 3/[(d_n/d) + 2\psi^{-0.5}]$, $K_2 = 10^{1.8148(-\text{Log}\psi)^{0.5743}}$ are two shape factors (d_n is the average between the minimum and the maximum axis, d is the equal volume sphere), and ψ is the particle sphericity ($\psi = 1$ for spheres). For calculating the sphericity is practical to use the concepts of operational and working sphericity, ψ_{work} introduced by Wadell (1933); Aschenbrenner (1956), which are based on the determination of the volume and of the three dimensions of a particle respectively:

$$\psi_{work} = 12.8 \frac{(P^2Q)^{1/3}}{1 + P(1+Q) + 6\sqrt{1 + P^2(1+Q^2)}} \quad (13)$$

with $P = S/I$, $Q = I/L$, where L is the longest particle dimension, I is the longest dimension perpendicular to L , and S is the dimension perpendicular to both L and I .

3. WILSON model (Walker et al., 1971; Wilson and Huang, 1979) using the interpolation suggested by Pfeiffer et al. (2005):

$$C_d = \begin{cases} \frac{24}{Re} \varphi^{-0.828} + 2\sqrt{1-\varphi} & Re \leq 10^2 \\ 1 - \frac{1 - C_d|_{Re=10^2}}{900} (10^3 - Re) & 10^2 \leq Re \leq 10^3 \\ 1 & Re \geq 10^3 \end{cases} \quad (14)$$

where $\varphi = (b+c)/2a$ is the particle aspect ratio ($a \geq b \geq c$ denote the particle semi-axes).

4. DELLINO model (Dellino et al., 2005):

$$v_s = 1.2605 \frac{\nu_a}{d} (Ar \xi^{1.6})^{0.5206} \quad (15)$$

where $Ar = gd^3(\rho_p - \rho_a)\rho_a/\mu_a^2$ is the Archimedes number, g the gravity acceleration, and ξ is a particle shape factor (sphericity to circularity ratio). It is recommended to not extrapolate this option for particle diameter beyond the range used in the experiments by Dellino et al. (2005).

Since for FALL3D-6.2 the primary particle shape factor is the sphericity ψ , for sake of simplicity, φ in (14) and ξ in (15) are calculated approximating particles as prolate ellipsoids (the same approximation is used for estimating d_n).

2.4 Meteorological variables

FALL3D-6.2 uses an off-line strategy, *i.e.* the meteorological variables are calculated independently by a different meteorological model or information, and interpolated to the spatial and temporal grid of FALL3D-6.2.

FALL3D-6.2 reads time-dependent meteorological data (wind field, air temperature, Monin-Obukhov length L , friction velocity u_* , and ABL height h) and topography from a database file in NetCDF format. This file can be created by an external utility program (SETDBS). This strategy allows FALL3D-6.2 to be used from micro- to meso-scale. In the PUB distribution version there are a few options to generate this database depending on the scale of application. For the possible choices see Section 4.2.

2.5 Source term

FALL3D-6.2 reads the time-dependent source term (mass released per unit time at each grid point) from an external file. This file can be generated by the SETSRC utility program, as *i*) a point source, *ii*) imposing a mushroom-like shape (Suzuki option) or *iii*) by using a model based on the Buoyant Plume Theory (BPT; see Section 4.4).

2.6 Particle aggregation

A subroutine describing particle aggregation in presence of water is being tested. Aggregation model and validation tests are presented in Costa et al. (2010); Folch et al. (2010).

3 Overview of the program FALL3D-6.2

FALL3D-6.2 needs the following input files:

- An input file where control parameters and options are specified (`filename.inp`). This file is read by FALL3D-6.2 and the utility programs.
- A database file in NetCDF format (see Appendix C) containing all meteorological data and the topography (`filename.dbs.nc`).
- A granulometry file specifying the characteristics of the particles emitted into the atmosphere (`filename.grn`).

- A source file specifying the discharge rates at the source points, typically along the eruptive column (`filename.src`).
- Optionally, a file specifying a list of points (`filename.pts`) where tracking of variables is performed (*e.g.* to compute ash arrival times, accumulation rates, etc).

The formats of the input files are described in Section 4. The FALL3D-6.2 package comes with a set of utility programs that can be used to generate the input files:

- The utility SETDBS can be used to generate the database file `filename.dbs.nc` created in accord to the parameters specified in the blocks TIME_UTC and GIRD of the input file `filename.inp`. This utility program reads meteo data from different sources and interpolates variables onto the FALL3D-6.2 computational grid. The time slice of the database must be equal or larger than the simulation time slice.
- The utility SETGRN can be used to generate the granulometry file `filename.grn` in accord to the parameters specified in the block GRANULOMETRY of the input file `filename.inp`. This program generates only Gaussian and Bi-Gaussian distributions. For other distributions the user must provide the granulometry file.
- The utility SETSRC can be used to generate the source file `filename.src` in accord to the parameters specified in the blocks TIME_UTC and SOURCE of the input file `filename.inp`.

The use of these utilities, although recommended, is not necessary if the user provides some of the necessary files directly.

Once a simulation is concluded, FALL3D-6.2 outputs the following files:

- The results file (`filename.res.nc`) in netCDF format. This file can be processed using several open-source programs (*e.g.* ncview, Panoply, ncl, etc.) to generate plots and animations.
- The log file (`filename.log`) contains information about the run, including summary of input data, error and warning messages, etc.
- The tracking points files (`filename.pts.*`) contain information about evolution of variables at the tracked points. There exist an output file for each point specified in the input file `filename.pts`.

4 The FALL3D-6.2 Input files

4.1 The control file `filename.inp`

The control input file is an ASCII file made up with a set of blocks that define all the computational and physical parameters needed by FALL3D-6.2 and the rest of utility programs. Each program reads only the necessary blocks of the file. Parameters within a block are listed one per record, in arbitrary order, and can optionally be followed by one or more blank spaces and a comment. Maximum allowed length is 256 characters including comments. A detailed description of each record is given below. Real numbers can be expressed following the FORTRAN notation (*e.g.*, $12e7 = 12 \times 10^7$).

- BLOCK TIME_UTC: Defines variables related to time.
- BLOCK GRID: Defines the characteristics of the FALL3D-6.2 computational mesh.
- BLOCK FALL3D: Defines the variables needed by FALL3D-6.2 program.
- BLOCK GRANULOMETRY: Defines the variables needed by SETGRN program.
- BLOCK SOURCE: Defines the variables needed by SETSRC program.
- BLOCK OUTPUT: Defines the FALL3D-6.2 strategy for dumping of results.

4.1.1 BLOCK TIME.UTC

This block of data defines variables related to time and is used by FALL3D-6.2 and by the utility programs SETDBS and SETSRC.

- **YEAR:** Database starting year.
- **MONTH:** Database starting month (1-12).
- **DAY:** Database starting day (1-31).
- **BEGIN_METEO_DATA_(HOURS_AFTER_00):** Time (in *h* after 0000UTC of the starting day) at which meteorological data starts in the database file.
- **TIME_STEP_METEO_DATA_(MIN):** Time step (in *min*) of the meteo data in the database file.
- **END_METEO_DATA_(HOURS_AFTER_00):** Time (in *h* after 0000UTC of the starting day) at which meteorological data ends in the database file. This time slice has to be larger than time slices defined by the records **ERUPTION_START_(HOURS_AFTER_00)** and **RUN_END_(HOURS_AFTER_00)**. If not, the program will stop.
- **ERUPTION_START_(HOURS_AFTER_00):** Eruption start hours (after 0000UTC of the day). These are *nt* values ($nt \geq 1$) that indicate the starting times of the different eruptive phases. Any type of transient behavior can be contemplated by adding a sufficient number of intervals. Eruptive conditions (plume height, MFR, etc.) are assumed constant during each phase (*i.e.* a quasi-steady approximation is used). The first value must be equal or larger than the value of the record **BEGIN_METEO_DATA_(HOURS_AFTER_00)**.
- **ERUPTION_END_(HOURS_AFTER_00) :** Eruption end hour (after 0000UTC of the starting day). If the SETSRC program is used to generate the source term, this is the time slice at which source term is switched off (*i.e.* the time at which the last eruptive phase ends).
- **RUN_END_(HOURS_AFTER_00):** Run end hour (after 0000UTC of the starting day). Must be equal or lower than the value of the record **END_METEO_DATA_(HOURS_AFTER_00)**. Note that, in general, a run should continue even when the source term is switched off (*i.e.* when the eruption has stopped) in order to allow the remaining airborne particles to sediment completely.

4.1.2 BLOCK GRID

This block of data defines the variables needed by SETDBS program to generate the FALL3D-6.2 grid. Note that time and spatial coverage of the database must include the FALL3D-6.2 simulation interval.

- **COORDINATES:** Defines the map projection. Possibilities are UTM or LON-LAT. Note that the UTM option can only be used if the domain is within a unique UTM zone. The use of the UTM coordinate system in large domains covering more than one UTM zone is not allowed (in this case, the LON-LAT option accounting for Earth's curvature must be used instead). The sub-blocks **UTM** or **LON_LAT** are read in each case respectively.
- **LONMIN:** Minimum longitude (in decimal degrees) of the domain (*i.e.* longitude corresponding to the bottom left corner). Only used in the LON-LAT option.
- **LONMAX:** Maximin longitude (in decimal degrees) of the domain (*i.e.* longitude corresponding to top right corner). Only used in the LON-LAT option.
- **LATMIN:** Minimum latitude (in decimal degrees) of the domain (*i.e.* latitude corresponding to bottom left corner). Only used in the LON-LAT option.
- **LATMAX:** Maximin latitude (in decimal degrees) of the domain (*i.e.* latitude corresponding to top right corner). Only used in the LON-LAT option.
- **LON_VENT:** Vent longitude. Only used in the LON-LAT option.
- **LAT_VENT:** Vent latitude. Only used in the LON-LAT option.

- **UTMZONE**: UTM zone code in format nnL (*e.g.* 33S). Only used in the UTM option.
- **XMIN**: minimum x -coordinate of the domain (bottom left corner). UTM coordinates must be given in m. Only used in the UTM option.
- **XMAX**: maximum x -coordinate of the domain (top right corner). UTM coordinates must be given in m. Only used in the UTM option.
- **YMIN**: minimum y -coordinate of the domain (bottom left corner). UTM coordinates must be given in m. Only used in the UTM option.
- **YMAX**: maximum y -coordinate of the domain (top right corner). UTM coordinates must be given in m. Only used in the UTM option.
- **X_VENT**: x -coordinate of the vent. UTM coordinates must be given in m. Only used in the UTM option.
- **Y_VENT**: y -coordinate of the vent. UTM coordinates must be given in m. Only used in the UTM option.
- **NX**: Number of grid nodes in the x -direction.
- **NY**: Number of grid nodes in the y -direction.
- **ZLAYER_(M)**: Array of heights (in m) of the vertical z -layers in terrain following coordinates. It is not necessary to specify the number of vertical layers since it is automatically calculated by the program. The vertical layers can be specified manually (as an array of values) or, for equally spaced vertical discretization, simply indicating the limits and the increment (*e.g.* FROM 0 TO 10000 INCREMENT 1000).

4.1.3 BLOCK FALL3D

This block of data defines the variables needed by FALL3D-6.2 program.

- **TERMINAL_VELOCITY_MODEL**: Type of terminal settling velocity model. Possibilities are ARASTOPOUR, GANSER, WILSON, and DELLINO.
- **VERTICAL_TURBULENCE_MODEL**: Type of model for vertical diffusion. Possibilities are CONSTANT or SIMILARITY.
- **VERTICAL_DIFFUSION_COEFFICIENT_(M2/S)**: Value of the diffusion coefficient (in m^2/s). Only used if **VERTICAL_TURBULENCE_MODEL** = CONSTANT
- **HORIZONTAL_TURBULENCE_MODEL**: Type of model for horizontal diffusion. Possibilities are CONSTANT, RAMS, or CMAQ.
- **HORIZONTAL_DIFFUSION_COEFFICIENT_(M2/S)**: Value of the diffusion coefficient (in m^2/s). Only used if **HORIZONTAL_TURBULENCE_MODEL** = CONSTANT.
- **RAMS_CS**: Value of C_S in the RAMS model (see eq. 6). Only used if **HORIZONTAL_TURBULENCE_MODEL** = RAMS.

4.1.4 BLOCK GRANULOMETRY

This block of data defines the variables needed by SETGRN program.

- **DISTRIBUTION**: Type of distribution. Possibilities are GAUSSIAN or BIGAUSSIAN.
- **NUMBER_OF_CLASSES**: Number of granulometric classes.
- **FI_MEAN**: Mean value of Φ (Gaussian distribution). For Bi-Gaussian distributions two values must be provided.

- **FI_DISP**: Value of σ (Gaussian distribution). For Bi-Gaussian distributions two values must be provided.
- **FI_RANGE**: Minimum and maximum values of Φ (Φ_{min} and Φ_{max} respectively).
- **DENSITY_RANGE**: Values of densities ρ_{min} and ρ_{max} (in kg/m^3) associated to Φ_{min} and Φ_{max} particles. Lineal interpolation is assumed. In particular, if $\rho_{min} = \rho_{max}$, density is constant for all classes.
- **SPHERICITY_RANGE**: Values of sphericity ψ_{min} and ψ_{max} associated to Φ_{min} and Φ_{max} particles. Lineal interpolation is assumed. In particular, if $\psi_{min} = \psi_{max}$, sphericity is constant for all classes.

4.1.5 BLOCK SOURCE

This block of data defines the variables needed by the SETSRC program. This program generates the source term (eruptive column) for each of the $nt \geq 1$ eruptive phases.

- **VENT_HEIGHT**: Height of the vent a.s.l. (in m).
- **SOURCE_TYPE**: Type of source distribution. Possibilities are POINT, SUZUKI or PLUME.

In the case **SOURCE_TYPE** = POINT only the sub-block POINT_SOURCE is used:

- **MASS_FLOW_RATE_(KGS)**: Array of values of the mass flow rate (in kg/s) for the nt eruptive phases. Alternatively, the user can use the word **estimate** and SETSRC automatically computes the MFR from the column heights based on empirical fits (Mastin et al., 2009). This is the typical situation during an eruption, when column height is likely to be the only observable available.
- **HEIGHT_ABOVE_VENT_(M)**: Array of heights of the plume (in m above the vent) for the nt eruptive phases. Note that the plume heights must be lower than the top of the computational domain, specified in the record ZLAYER_(M) of the GRID block. If not, the program will stop.

In the case **SOURCE_TYPE** = SUZUKI only the sub-block SUZUKI_SOURCE is used:

- **MASS_FLOW_RATE_(KGS)**: Array of values of the mass flow rate (in kg/s) for the nt eruptive phases. Alternatively, the user can use the word **estimate** and SETSRC automatically computes the MFR from the column heights based on empirical fits (Mastin et al., 2009). This is the typical situation during an eruption, when column height is likely to be the only observable available.
- **HEIGHT_ABOVE_VENT_(M)**: Array of heights of the plume (in m above the vent) for the nt eruptive phases. Note that the plume heights must be lower than the top of the computational domain, specified in the record ZLAYER_(M) of the GRID block. If not, the program will stop.
- **A**: Array of values of the parameter A in the Suzuki distribution for the nt eruptive phases (Pfeiffer et al., 2005).
- **L**: Array of values of the parameter λ in the Suzuki distribution for the nt eruptive phases (Pfeiffer et al., 2005).

In the case **SOURCE_TYPE** = PLUME only the sub-block PLUME_SOURCE is used:

- **SOLVE_PLUME_FOR**: Possibilities are MFR or HEIGHT. In the first case SETSRC solves for the mass flow rate given the column height, whereas in the second does the opposite. Since the plume equations use the mass flow rate as an input, the first option requires an iterative procedure.
- **MFR_SEARCH_RANGE**: Two values n and m such that 10^n and 10^m specify the range of MFR values admitted in the iterative solving procedure (*i.e.* it is assumed that $10^n \leq MFR \leq 10^m$). Only used if SOLVE_PLUME_FOR=MFR.
- **MASS_FLOW_RATE_(KGS)**: Values of the mass flow rate (in kg/s) for the nt eruptive phases. Only used if SOLVE_PLUME_FOR=HEIGHT.

- **HEIGHT_ABOVE_VENT_(M)**: Heights of the plume (in m above the vent) for the *nt* eruptive phases. Note that the plume heights must be lower than the top of the computational domain, specified in the record **ZLAYER_(M)** of the **GRID** block. Only used if **SOLVE_PLUME_FOR=MFR**.
- **EXIT_VELOCITY_(MS)**: Values of the magma exit velocity (in m/s) at the vent for the *nt* eruptive phases.
- **EXIT_TEMPERATURE_(K)**: Values of the magma exit temperature (in K) at the vent for the *nt* eruptive phases.
- **EXIT_VOLATILE_FRACTION_(IN%)**: Values of the magma volatile at the vent for the *nt* eruptive phases in weight percent.

4.1.6 BLOCK OUTPUT

This block of data defines the output strategy of the FALL3D-6.2 program.

- **POSTPROCESS_TIME_INTERVAL_(HOURS)**: Postprocess time interval in hours.
- **POSTPROCESS_3D_VARIABLES**: Possibilities are YES or NO. If YES, FALL3D-6.2 writes 3D concentration in the output file `filename.res.nc`. If NO, only 2D variables are written to the output file.
- **POSTPROCESS_CLASSES**: Possibilities are YES or NO. If YES, FALL3D-6.2 writes results for all the classes. If NO, only total results are written.
- **TRACK_POINTS**: Possibilities are YES or NO. If YES, FALL3D-6.2 writes the tracking points files.

4.2 The database file `filename.dbs.nc`

This file written in NetCDF format (see Appendix C) contains time-dependent meteorological data (wind field, air temperature and density, humidity, etc) needed by FALL3D-6.2. The file can be created by the external utility program SETDBS, which reads meteo data and interpolates to the FALL3D-6.2 space-time domain. There are several options to generate this database depending on the scale of application. This strategy allows FALL3D-6.2 to be used from micro- to meso-scale. The possible choices are described below.

- The simplest option consists of using a horizontally uniform wind derived from a vertical profile, typically obtained from sounding measurements or from indirect reconstructions. The vertical profile needs to be specified in the file `filename.profile` in the format described in the Appendix A. In this case, in addition to the profile `filename.profile` it is also necessary to furnish a topography file `filename.top` in GRD format (see Appendix B).
- The second choice (**CALMET** option) uses data derived from the output of the meteorological diagnostic model CALMET (Scire et al., 2000). This option is used for assimilating and interpolating short-term forecasts (or re-analysis) from Mesoscale Meteorological Prognostic Models (MMPM) to a finer scale. In this case only the UTM coordinate system can be used. Note that the output of CALMET is a binary file that depends on the architecture of the machine were it was generated. Moreover note that this option is compatible only with a CALMET output time step equal to an hour (*i.e.*, `nsecdt=3600`).
- The third choice (**NCEP-1** option) uses data from NCEP re-analysis 1 (see: <http://www.esrl.noaa.gov/psd/data/gridded/data.ncep.reanalysis.html>).
- Other choices (available only in the PROF version) contemplate several global and mesoscale meteorological models such as ARPA-SIM, ETA, GFS, or NMMb.

nc
diam(1) rho(1) sphe(1) fc(1)
...
diam(nc) rho(nc) sphe(1) fc(nc)

Table 2: Format of the granulometry file `filename.grn`.

4.3 The granulometry file `filename.grn`

The granulometry file is an ASCII file containing the definition of the particle classes (a class is characterized by particle size, density and sphericity). This file can be created by the utility program `SETGRN`. Note that `SETGRN` only generates distributions which are Gaussian or bi-Gaussian in ψ (log-normal in d) and linear in ρ and ψ . `FALL3D-6.2` can handle more general distributions but, in this case, the granulometry file `filename.grn` has to be defined directly by the user. The file format is described in Table 2 and the meaning of the used symbols is the following:

- **nc**: Number of particle classes.
- **diam**: Class diameter (in mm).
- **rho**: Class density (in kg/m³).
- **sphe**: Class sphericity.
- **fc**: Class mass fraction (0-1). It must verify that $\sum \mathbf{fc} = 1$.

4.4 The source file `filename.src`

The source file `filename.src` is an ASCII file containing the definition of the source term. The source can be defined for different time phases during which source values are kept constant. The number, position and values (*i.e.* Mass Flow Rate) of the source points can vary from one time slice to another and cannot overlap. There is no restriction on the number and duration of the time slices. It allows, in practise, to discretize any kind of source term. This file can be defined directly by the user, in the format described in Table 3, or created by using the utility program `SETSRC`. In the last case, the `filename.src` is created in accord to the parameter specified in the `SOURCE` block of the `filename.inp`. The options to be chosen in the `filename.inp` are *i*) a point source, *ii*) a mushroom-like shape (Suzuki option) or *iii*) an eruption column model based on the Buoyant Plume Theory (BPT). The format of the file `filename.src` is described in Table 3 and the meaning of the used symbols is the following:

- **itime1**: Starting time of the time slice (in sec after 00UTC of the starting day).
- **itime2**: End time of the time slice (in sec after 00UTC of the eruption starting day).
- **nsrc**: Number of source points (can vary from one interval to another depending on the column height).
- **nc**: Number of particle classes.
- **MFR**: Mass flow rate (in kg/s).
- **x**: Longitude or x -coordinate of the source *isrc*.
- **y**: Latitude or y -coordinate of the source *isrc*.
- **z**: z -coordinate of the source *isrc* above ground level (a.g.l.) (in m).
- **src**: Mass flow rate (in kg/s) of each granulometric class for this point source. It must be verified that $\sum \sum \mathbf{src}(isrc, ic) = MFR$.

```

itime1 itime2
nsrc nc
MFR
x y z src(1,1) ... src(1,nc)
...
x y z src(nsrc,1) ... src(nsrc,nc)

```

Table 3: Format of the source file `filename.src`. Repeat this block for each time slice

4.5 The points file `filename.pts`

This file contains the names (identifiers) and coordinates of the points to be tracked. It is used only when the record `TRACK_POINTS` in the input file `filename.inp` is set to `YES`. The format of the file `filename.pts` consists of lines (one line per point) with three columns specifying the point name, the point longitude (or x -coordinate if UTM coordinates are used), and the point latitude (or y -coordinate if UTM coordinates are used). There is no limit on the number of points.

5 Program Setup

5.1 Installation

To install FALL3D-6.2 and the utility programs uncompress and untar the file `Fall3d-6.2.PUB.tar.gz`. It will create the folder structure shown in Table 4. The package contains the source codes, scripts, documentation, and a run examples.

- For Unix/Linux/Mac X OS it is necessary to have a FORTRAN compiler and the NetCDF library (<http://www.unidata.ucar.edu/software/netcdf/>) already installed (version 3.6 or later). It is mandatory to compile FALL3D-6.2 using the same FORTRAN compiler that has been used to compile the NetCDF library.

5.1.1 Unix/Linux/Mac X OS

For Unix/Linux/Mac X OS the package comes with an automatic installation script. Proceed as follows:

1. Enter the directory `Install` and edit the `Install` script file. Set up the variables `HOMEFALL3D` (FALL3D-6.2 home directory path), `Lib_netcdf` (path of the NetCDF library in your computer), and `COMPILER` (name of the FORTRAN compiler).
NOTE: Automatic installation is possible for the following standard compilers: `gfortran`, `ifort`, `f90`, `xlf90`. If you want to compile using a different compiler it is necessary to modify the Makefiles and the Scripts manually.
2. Run the `Install` script. This will compile FALL3D-6.2 and the utility programs, modify the scripts introducing your FALL3D-6.2 path and check the installation process.
3. Optionally, create an alias to the `Script-manager` file located in the folder `Scripts`. This script allows for launching FALL3D-6.2 and the utility programs directly from the command line.

5.1.2 Windows OS

Not yet available for the current versions.

5.2 Execution

To create a new run named `problemname` simply create a new directory `problemname` in the folder `Runs`, copy the control input file from the example run (rename it as `problemname.inp`), and modify it depending on your needs.

Level 1	Level 2	Level 3	Description
Fall3d	Documents		Contains the manual.
	Install		Contains installation scripts.
	Runs	Run-name	Contains the examples, one folder each.
	Scripts		Contains the script files.
	Sources_ser		FALL3D-6.2 sources (PUB version).
	Utilities	LibMaster	Master library.
		SetDbs	SETDBS utility program.
		SetGrn	SETGRN utility program.
		SetSrc	SETSRC utility program.

Table 4: Default structure of FALL3D-6.2 folders.

5.2.1 Unix/Linux/Mac X OS

FALL3D-6.2 and the utility programs can be launched using the `Script-manager` with a series of arguments. It is recommended to use an alias for this script that can be called directly from any location (in the following it is assumed that the alias is `Launch`). From any location:

- Type `Launch SetGrn problemname` to run the SETGRN utility program for `problemname`.
- Type `Launch SetDbs problemname meteo` to run the SETDBS utility program for `problemname`. Here `meteo` is one of the following: `profile/calmet62/ncep1`.
- Type `Launch SetSrc problemname` to run the SETSRC utility program for `problemname`.
- Type `Launch Pub problemname` to run the PUB version of FALL3D-6.2 for `problemname`.

5.2.2 Windows OS

Not yet available for the current versions.

Appendices

Appendix A: Format of the meteo profile file (`filename.profile`)

For the profile option, the utility `SetDbs` needs an ASCII file containing the definition of the vertical wind profile and a topography file of the domain in GRD format (see Appendix B). In this case wind velocities are assumed constant on all the domain in a terrain-following coordinate system. The remaining variables are assumed with the values of the Standard Atmosphere. The format of the profile file (`filename.profile`) is described in Table 5 and the meaning of the used symbols is the following:

- `pcoord`: Coordinates where the profile was measured; either as UTM or lon-lat coordinates.
- `pdate`: Starting time when the profile was measured; the format of the date is `yyyymmdd`, *i.e.* year, month, day.
- `itime1`: Initial time in sec after the starting time `pdate` of validity of the meteo data contained in the following `nz` layers.
- `itime2`: Final time in sec after the starting time `pdate` of validity of the meteo data contained in the following `nz` layers.
- `nz`: Number of the database vertical layers.
- `z`: Vertical coordinate of the layer (in m a.s.l.).
- `ux`: wind x -velocity (in m/s).
- `uy`: wind y -velocity (in m/s).

```

pcoord
pdate
itime1 itime2
nz
z(1) ux(1) ux(1) T(1)
...
z(nz) ux(nz) ux(nz) T(nz)
itime3 itime4
...

```

Table 5: Format of the meteo data file `filename.profile.dat` for the PROFILE case. Repeat this block for each meteo time increment.

Appendix B: The GRD format

The structure of a GRD format file is described in Table 6 and the meaning of the used symbols is the following:

- **NX** : Number of grid points along x -direction.
- **NY** : Number of grid points along y -direction.
- **X0** : x -coordinate (UTM in m) of the grid bottom left corner.
- **XF** : x -coordinate (UTM in m) of the grid top right corner point.
- **Y0** : y -coordinate (UTM in m) of the grid bottom left corner point.
- **YF** : y -coordinate (UTM in m) of the grid top right corner point.
- **VAL** : Value at each grid point. It consists of an array of $NX \times NY$ values stored starting from the bottom-left corner and moving towards right then up towards the top-right corner.

NX	NY		
X0	XF		
Y0	YF		
MAX(v)	MIN(v)		
VAL(i,1)	i=1:NX
...	
VAL(i,j)	i=1:NX
...	
VAL(i,NY)	i=1:NX

Table 6: Format of a GRD file `filename.grd`.

Appendix C: The NetCDF format

NetCDF (network Common Data Form) is a set of software libraries and machine-independent data formats that support the creation, access, and sharing of array-oriented scientific data (available at: <http://www.unidata.ucar.edu/software/netcdf/>). FALL3D-6.2 uses the standard NetCDF format for both database input file (`filename.dbs.nc`) and results output file (`filename.res.nc`). Only the PROF version comes with the utility to view, manipulate or transform NetCDF files. However, there is a good number of open-source codes to do so. For example:

- `ncview` and `ncdump` (http://opendap.org/download/nc_clients.html).
- `Panoply` (<http://www.giss.nasa.gov/tools/panoply/>).

- GrADS (<http://www.iges.org/grads/>).
- NCL, the NCAR Command Language (<http://www.ncl.ucar.edu/>).

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